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# Best Linear Unbiased and Invariant Reconstructors for the Past Records

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**Abstract.** The best linear unbiased reconstructor and the best linear invariant reconstructor of the past upper records based on observed upper record values for the location–scale family of distributions are derived. The results are obtained in detail for two-parameter exponential distribution. A comparison study is performed in terms of mean squared reconstruction error and Pitman's measure of closeness criteria. Finally, a real example is given to illustrate the proposed procedures.

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#### 1. Introduction

Let  $\{X_i, i > 1\}$  be a sequence of independent and identically distributed (iid) continuous random variables. An observation  $X_j$  is called an upper record value if its value exceeds that of all previous observations, i. e.,  $X_j$  is an upper record value if  $X_j > X_i$  for every i < j. These type of data are of great importance in several real-life problems involving weather, industry, economic and sport data. For more details and applications of record values, see for example, Arnold *et al.* [4]. Absanullah and Shakil [3] obtained some characterization results of Rayleigh distribution based on conditional expectation of record values.

The prediction problem of the future records based on past records have been extensively studied in both frequentist and Bayesian approaches; Awad and Raqab [6] considered the prediction problem of the future *n*th record value based on the first m (m < n) observed record values from one-parameter exponential distribution. Raqab [17] established different point predictors and prediction intervals for future records based on observed record values from two-parameter exponential distribution. Raqab and Balakrishnan [18] discussed the problem of prediction of the future record values in nonparametric settings. Ahmadi and Doosparast [2] investigated Bayesian estimation and prediction for some life distributions based on upper record values. MirMostafaee and Ahmadi [16] obtained several point

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predictors such as the best unbiased predictor, the best invariant predictor and maximum likelihood predictor of future order statistics on the basis of observed record values from two-parameter exponential distribution.

Recently, some works have been done on the reconstruction problem. Klimczak and Rychlik [14] obtained upper bounds for the expectations of increments of past order statistics and records based on observed order statistics and records, respectively. Balakrishnan et al. [10] investigated the problem of reconstructing past upper records from the known values of future records in exponential and Pareto distributions. Razmkhah et al. [19] obtained some point and interval reconstructors for the missing order statistics in exponential distribution. Asgharzadeh et al. [5] studied the reconstruction of the past failure times based on left-censored samples for the proportional reversed hazard rate models. The main goal of this paper is to investigate the best linear unbiased reconstructor (BLUR) and the best linear invariant reconstructor (BLIR) based on observed future record data from two-parameter exponential distribution. We also intend to present a comparison study of common reconstructors of past records from exponential distribution. With this in mind, we consider two criteria: mean squared reconstruction error (MSRE) and Pitman's measure of closeness (PMC). In recent years, many problems involving ordered data and Pitman closeness have been investigated. Pitman closeness of records to population quantiles was explored in Ahmadi and Balakrishnan [1]. Some interesting results regarding the sample median and Pitman closeness were provided by Iliopoulous and Balakrishnan [13]. Based on a Type-II right censored sample from one-parameter exponential distribution, Balakrishnan et al. [9] compared the best linear unbiased predictor (BLUP) and the best linear invariant predictor (BLIP) of the censored order statistics in the one-sample case in terms of PMC.

The rest of this paper is organized as follows. Section 2 contains some preliminaries. In Sections 3 and 4, the BLUR and BLIR of the past upper records from two-parameter exponential distribution are investigated, respectively. These two reconstructors and two others where obtained by Balakrishnan *et al.* [10] are compared in terms of MSRE and PMC criteria in Section 5. A real example is given in Section 6.

#### 2. Preliminaries

Let  $R_1, ..., R_n$  be the first *n* upper record values coming from a *X*-sequence of iid continuous random variables with cumulative distribution function (cdf)  $F(x; \theta)$  and probability density function (pdf)  $f(x; \theta)$ , where  $\theta \in \Theta$  may be a vector of parameters and  $\Theta$  is the parameter space. Suppose we have failed to observe the first *m* records, namely, we have observed  $\mathbf{R} = (R_{m+1}, R_{m+2}, ..., R_n), m+1 < n$ . In this case, the joint pdf of **R** can be expressed as follows (see, for example, Arnold *et al.* [4, pp. 10–11])

(2.1) 
$$f_{\mathbf{R}}(\mathbf{r}; \theta) = \frac{\left(H(r_{m+1}; \theta)\right)^m}{m!} f(r_{m+1}; \theta) \prod_{i=m+1}^{n-1} \frac{f(r_{i+1}; \theta)}{\bar{F}(r_i; \theta)}, r_{m+1} < \ldots < r_n,$$

where  $\bar{F}(x;\theta) = 1 - F(x;\theta)$  is the survival function of X,  $H(x;\theta) = -\log \bar{F}(x;\theta)$ ,  $\mathbf{r} = (r_{m+1}, \ldots, r_n)$  is the observed value of  $\mathbf{R}$ . From (2.1) and using the Markovian property of record values, it can be shown that the conditional pdf of  $R_l$  given the observed data set  $\mathbf{R}$  is just the conditional pdf of  $R_l$  given  $R_{m+1}$  and is given by

$$f_{R_l|\mathbf{R}}(r_l|\mathbf{r};\boldsymbol{\theta}) = f_{R_l|R_{m+1}}(r_l|r_{m+1};\boldsymbol{\theta})$$

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(2.2) 
$$= \frac{(H(r_l; \theta))^{l-1} (H(r_{m+1}; \theta) - H(r_l; \theta))^{m-l}}{B(l, m-l+1) (H(r_{m+1}; \theta))^m} \frac{f(r_l; \theta)}{\bar{F}(r_l; \theta)}, r_l < r_{m+1},$$

where B(.,.) is the complete beta function. From (2.2), the unbiased reconstructor of  $R_l$ , l = 1, ..., m, is the conditional mean of  $R_l$  given  $R_{m+1}$ , i.e.,

(2.3) 
$$R_l^* = E(R_l | \mathbf{R}) = E(R_l | R_{m+1}),$$

and hence it depends only on  $R_{m+1}$ . If the parameters in (2.3) are unknown, they have to be estimated in this conditional expectation.

A random variable X is said to have a two-parameter exponential distribution, denoted by  $\text{Exp}(\mu, \sigma)$ , if its cdf is

(2.4) 
$$F(x;\theta) = 1 - e^{-\frac{x-\mu}{\sigma}}, x \ge \mu, \ \sigma > 0,$$

where  $\theta = (\mu, \sigma)$ ,  $\mu$  and  $\sigma$  are the location and scale parameters, respectively. The exponential distribution is a very commonly used distribution in reliability engineering and survival analysis. We refer the reader to Balakrishnan and Basu [7] for some research based on exponential distribution. In what follows, we shall study the behavior of reconstructors of the past records for  $\text{Exp}(\mu, \sigma)$  in detail. The first result for the case of  $\text{Exp}(\mu, \sigma)$  is that the unbiased reconstructor of  $R_l$  based on **R** is given by

(2.5) 
$$R_l^*(\mu) = \mu + \frac{l}{m+1}(R_{m+1} - \mu),$$

it does not depend on the scale parameter. When  $\mu$  is unknown, we can substitute it with its estimator based on  $\mathbf{R} = (R_{m+1}, R_{m+2}, ..., R_n)$ .

## 3. The best linear unbiased reconstructor

In this section, first we focus our attention to derive the BLUR in general for the locationscale families. Let  $\mathbf{X} = (X_{m+1}, ..., X_n)'$  be an observed (n-m)-dimensional random vector from the cdf  $F(x; \mu, \sigma) = F_0((x - \mu)/\sigma)$ , where  $\mu$  and  $\sigma$  are the location and scale parameters, respectively, and  $F_0$  does not depend on  $\mu$  and  $\sigma$ . Also, let Y be an unknown past observation from the same distribution. Reconstructing the outcome Y based on the observed value of  $\mathbf{X}$  is the main goal of this section. Toward this end, let us assume that  $a_0 = E((Y - \mu)/\sigma), a_i = E((X_i - \mu)/\sigma), w_i = Cov((Y - \mu)/\sigma, (X_i - \mu)/\sigma)$  and  $v_{i,j} = Cov((X_i - \mu)/\sigma, (X_j - \mu)/\sigma)$ , for i, j = m + 1, ..., n. Denote by 1 the vector of 1 vs,  $\mathbf{a} = (a_{m+1}, ..., a_n)^T$ ,  $\mathbf{w}^T = (w_{m+1}, ..., w_n)$  and  $\mathbf{V} = (v_{i,j}), i, j = m + 1, ..., n$ . To obtain the BLUR of Y based on  $\mathbf{X}$ , we will have to minimize the variance of  $(Y - \mathbf{cX}^T)$  subject to the condition of unbiasedness where  $\mathbf{c} = (c_{m+1}, ..., c_n)$  is an arbitrary vector of real values. Then, following along the same lines as in the prediction problem, the BLUR of Y is given by (see, Goldberger [12] in the context of prediction)

(3.1) 
$$\hat{Y} = \hat{\mu}(\mathbf{X}) + \hat{\sigma}(\mathbf{X})a_0 + \mathbf{w}^T \mathbf{V}^{-1}[\mathbf{X}^T - \hat{\mu}(\mathbf{X})\mathbf{1} - \hat{\sigma}(\mathbf{X})\mathbf{a}],$$

where  $\hat{\mu}(\mathbf{X})$  and  $\hat{\sigma}(\mathbf{X})$  are the BLUEs of  $\mu$  and  $\sigma$  on the basis of  $\mathbf{X}$ , respectively, and  $\mathbf{V}^{-1}$  is the inverse matrix of  $\mathbf{V}$ .

Now, assume that the observed data are the upper record values from the two-parameter exponential distribution. Using (3.1), the BLUR of  $R_l$  based on **R** is given by

(3.2) 
$$\hat{R}_l = \hat{\mu} + l\hat{\sigma} + \omega^T \Sigma^{-1} [\mathbf{R}^T - \hat{\mu} \mathbf{1} - \hat{\sigma} \alpha],$$

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where  $\hat{\mu}$  and  $\hat{\sigma}$  are the BLUEs of  $\mu$  and  $\sigma$  based on **R**, respectively,  $\alpha$  is the vector of means of the corresponding record values taken from Exp(0,1),  $\sigma^2 \Sigma$  is the covariance matrix of **R** and

$$\omega^{T} = 1/\sigma^{2}(Cov(R_{m+1},R_{l}),Cov(R_{m+2},R_{l}),...,Cov(R_{n},R_{l}))$$

By doing some algebraic calculations it can be shown that

$$\boldsymbol{\omega}^T = (l, l, ..., l)$$

and

(3.3) 
$$\omega^T \Sigma^{-1} = (\frac{l}{m+1}, 0, 0, ..., 0).$$

Substituting (3.3) into (3.2), we get

$$\hat{R}_l = \hat{\mu} + l\hat{\sigma}.$$

Furthermore, the BLUEs of  $\mu$  and  $\sigma$  on the basis of **R** are given by

(3.5) 
$$\hat{\mu} = \frac{nR_{m+1} - (m+1)R_n}{n - (m+1)} \text{ and } \hat{\sigma} = \frac{R_n - R_{m+1}}{n - (m+1)},$$

respectively (see, Arnold *et al.* [4]). Therefore, by substituting (3.5) in (3.4), it is deduced that

(3.6) 
$$\hat{R}_l = \frac{(n-l)R_{m+1} - ((m+1)-l)R_n}{n - (m+1)}$$

is the BLUR of  $R_l$  on the basis of the data set **R**. The MSRE of  $\hat{R}_l$  is computed to be

(3.7) 
$$MSRE(\hat{R}_l) = E[\hat{R}_l - R_l]^2 = \sigma^2 \frac{(n-l)((m+1)-l)}{n-(m+1)}$$

It is easy to show that  $\text{MSRE}(\hat{R}_l)$  is increasing in *m* and decreasing with respect to *n* and *l*, if the other arguments are fixed. Moreover,  $\hat{R}_l$  and  $\text{MSRE}(\hat{R}_l)$  converge to  $R_{m+1} - \hat{\sigma}$  and  $\sigma^2(n-m)/(n-(m+1))$ , respectively, when *l* tends to *m*.

**Remark 3.1.** It may be noted that by substituting the BLUE of  $\mu$  into the unbiased reconstructor of  $R_l$  in (2.5), the same result will be obtained for  $\hat{R}_l$  as in (3.6), i.e.,  $R_l^*(\hat{\mu}) = \hat{R}_l$ .

## 4. The best linear invariant reconstructor

Same as in previous section, first we present the results for BLIR in general for location–scale families. Let  $\mathbf{X} = (X_{m+1}, ..., X_n)'$  and *Y* be as defined in Section 3. Then, follow the results of Mann [15], in the context of prediction, the BLIR of *Y* based on the data set **X** can be derived as

(4.1) 
$$\widetilde{Y} = \widehat{Y} - \left(\frac{c_1}{1+c_2}\right)\widehat{\sigma}(\mathbf{X}),$$

where  $\hat{Y}$  is the BLUR of Y,  $c_2 = \sigma^{-2} Var(\hat{\sigma}(\mathbf{X}))$  and

$$c_1 = 1/\sigma^2 Cov(\hat{\boldsymbol{\sigma}}(\mathbf{X}), (1 - \mathbf{w}^T \mathbf{V}^{-1} \mathbf{1})\hat{\boldsymbol{\mu}}(\mathbf{X}) + (a_0 - \mathbf{w}^T \mathbf{V}^{-1} \mathbf{a})\hat{\boldsymbol{\sigma}}(\mathbf{X})),$$

with **V**, **w**, **a**,  $\hat{\mu}(\mathbf{X})$ ,  $\hat{\sigma}(\mathbf{X})$  and  $a_0$  being as defined in Section 3. When the upper record statistics  $\mathbf{R} = (R_{m+1}, R_{m+2}, ..., R_n)$  from the Exp $(\mu, \sigma)$  distribution are observed, using (4.1), the BLIR of  $R_l$  on the basis of **R** is as follows

(4.2) 
$$\widetilde{R}_l = \hat{R}_l - \left(\frac{V_1}{1+V_2}\right)\hat{\sigma},$$

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where  $V_2 = \sigma^{-2} Var(\hat{\sigma})$  and  $V_1 = \sigma^{-2} [Cov(\hat{\sigma}, (1 - \omega^T \Sigma^{-1} \mathbf{1})\hat{\mu} + (l - \omega^T \Sigma^{-1} \alpha)\hat{\sigma})]$  in which  $\omega, \Sigma^{-1}$  and  $\alpha$  are as defined in Section 3,  $\hat{\mu}$  and  $\hat{\sigma}$  are the BLUEs of  $\mu$  and  $\sigma$ , respectively, and  $\hat{R}_l$  is the BLUR of  $R_l$ . By performing some algebraic calculations, we obtain

(4.3) 
$$\widetilde{R}_{l} = \frac{(n+1-l)R_{m+1} - (m+1-l)R_{n}}{n-m}.$$

Moreover, it can be shown that the MSRE of  $\widetilde{R}_l$  is given by

(4.4) 
$$\operatorname{MSRE}(\widetilde{R}_l) = E[\widetilde{R}_l - R_l]^2 = \sigma^2 \frac{(n+1-l)(m+1-l)}{n-m},$$

which is increasing with respect to *m* and decreasing with respect to *n* and *l*, if the other arguments are fixed. Also,  $\tilde{R}_l$  and  $MSRE(\tilde{R}_l)$  converge to  $R_{m+1} - \tilde{\sigma}$  and  $\sigma^2(1 + 1/(n-m))$ , respectively, when *l* tends to *m*.

**Remark 4.1.** It may be noted that the BLIEs of  $\mu$  and  $\sigma$  for a two-parameter exponential distribution are given by (see, Arnold *et al.* [4])

(4.5) 
$$\widetilde{\mu} = \frac{(n+1)R_{m+1} - (m+1)R_n}{n-m} \text{ and } \widetilde{\sigma} = \frac{R_n - R_{m+1}}{n-m}$$

respectively. By plugging the BLIE of  $\mu$  into (2.5), we arrive at the same result for  $\widetilde{R}_l$  as given in (4.3), i.e.,  $R_l^*(\widetilde{\mu}) = \widetilde{R}_l$ .

## 5. Comparison results

In Sections 3 and 4, we obtained two reconstructors (BLUR and BLIR) for  $R_l$  based on **R** from a two-parameter exponential distribution. Balakrishnan *et al.* [10] also considered the same plan and obtained two other reconstructors for  $R_l$  namely maximum likelihood reconstructor (MLR), denoted by  $\hat{R}_{l,M}$ , and conditional median reconstructor (CMR), denoted by  $\hat{R}_{l,C}$ . For a comparison study, let us restate their results here. The MLR of  $R_l$  is

(5.1) 
$$\hat{R}_{l,M} = \frac{(n-l+1)R_{m+1} - (m-l)R_n}{n-m+1}$$

with

(5.2) 
$$MSRE(\hat{R}_{l,M}) = \sigma^2 \left\{ \frac{(n+1-l)(m+1-l)}{n-m+1} + \frac{(n+1+m-2l)(n+1-l)}{(n-m+1)^2} \right\}.$$

The CMR of  $R_l$  is

(5.3) 
$$\hat{R}_{l,C} = \frac{m(R_n - R_{m+1})M_{l,m} + nR_{m+1} - mR_n}{n - m}$$

with

(5.4)  

$$MSRE(\hat{R}_{l,C}) = \sigma^{2} \left\{ (m+1-l)(m+2-l) + \left(\frac{m(1-M_{l,m})}{n-m}\right)^{2}(n-m-1) \times (n-m) - 2\left(\frac{m(1-M_{l,m})}{n-m}\right)(m+1-l)(n-m-1) \right\},$$

where  $M_{l,m}$  stands for the median of beta distribution with parameters l and m-l+1. In this section, we intend to compare the four reconstructors based on MSRE and PMC criteria.

## 5.1. Comparison based on MSRE

Using (3.7) and (4.4), we have

(5.5) 
$$\operatorname{MSRE}(\widetilde{R}_l) = \operatorname{MSRE}(\widehat{R}_l) - \sigma^2 \frac{\left((m+1)-l\right)^2}{(n-(m+1))(n-m)}.$$

From (5.5), it is obvious that BLIR is better than BLUR in the sense of MSRE. Moreover, from (4.4) and (5.2), we have

(5.6) 
$$\mathrm{MSRE}(\widetilde{R}_l) = \mathrm{MSRE}(\widehat{R}_{l,M}) - \sigma^2 \left(\frac{n+1-l}{n-m+1}\right)^2 (1-\frac{1}{n-m}),$$

which implies that BLIR is better than MLR. Furthermore, using (3.7) and (5.2), we get

(5.7)

$$MSRE(\hat{R}_l) = MSRE(\hat{R}_{l,M}) - \sigma^2 \frac{n+1+m-2l}{(n-m+1)^2(n-m-1)} \times \left( (n-m)^2 - (n+m-2(l-1)) \right).$$

Using (5.7) it is deduced that BLUR is better than MLR when  $(n-m)^2 > (n+m-2(l-1))$ .

In order to complete the comparison study, we have computed the numerical values of MSREs of four reconstructors for n = 10 and all selected values of l and m. The results are presented in Table 1.

Table 1. Numerical values of MSREs for n = 10.

|   |                     |        |        |        | m       |         |         |        |
|---|---------------------|--------|--------|--------|---------|---------|---------|--------|
| l | 1                   | 2      | 3      | 4      | 5       | 6       | 7       | 8      |
| 1 | $2.0000^{1}$        | 3.5802 | 5.6250 | 8.3673 | 12.2222 | 18.0000 | 27.5000 | 45.555 |
|   | 1.3333 <sup>2</sup> | 2.6490 | 4.6140 | 7.0042 | 10.3351 | 15.3214 | 23.6237 | 40.220 |
|   | $1.1250^{3}$        | 2.5714 | 4.5000 | 7.2000 | 11.2500 | 18.0000 | 31.5000 | 72.000 |
|   | 1.1111 <sup>4</sup> | 2.5000 | 4.2857 | 6.6667 | 10.0000 | 15.0000 | 23.3333 | 40.000 |
| 2 |                     | 2.0000 | 3.6562 | 5.8776 | 9.0000  | 13.6800 | 21.3750 | 36.000 |
|   |                     | 1.1250 | 2.7857 | 4.7456 | 7.4591  | 11.5082 | 18.2395 | 31.685 |
|   |                     | 1.1429 | 2.6667 | 4.8000 | 8.0000  | 13.3333 | 24.0000 | 56.000 |
|   |                     | 1.1250 | 2.5714 | 4.5000 | 7.2000  | 11.2500 | 18.0000 | 31.500 |
| 3 |                     |        | 2.0000 | 3.7551 | 6.2222  | 9.9200  | 16.0000 | 27.555 |
|   |                     |        | 1.2673 | 2.8408 | 5.0000  | 8.2094  | 13.5341 | 24.159 |
|   |                     |        | 1.1667 | 2.8000 | 5.2500  | 9.3333  | 17.5000 | 42.000 |
|   |                     |        | 1.1429 | 2.6667 | 4.8000  | 8.0000  | 13.3333 | 24.000 |
| 4 |                     |        |        | 2.0000 | 3.8889  | 6.7200  | 11.3750 | 20.222 |
|   |                     |        |        | 1.2768 | 2.9486  | 5.4168  | 9.5000  | 17.635 |
|   |                     |        |        | 1.2000 | 3.0000  | 6.0000  | 12.0000 | 30.000 |
|   |                     |        |        | 1.1667 | 2.8000  | 5.2500  | 9.3333  | 17.500 |
| 5 |                     |        |        |        | 2.0000  | 4.0800  | 7.5000  | 14.000 |
|   |                     |        |        |        | 1.2995  | 3.1281  | 6.1357  | 12.114 |
|   |                     |        |        |        | 1.2500  | 3.3333  | 7.5000  | 20.000 |
|   |                     |        |        |        | 1.2000  | 3.0000  | 6.0000  | 12.000 |
| 6 |                     |        |        |        |         | 2.0000  | 4.3750  | 8.8889 |
|   |                     |        |        |        |         | 1.3395  | 3.4403  | 7.5950 |
|   |                     |        |        |        |         | 1.3333  | 4.0000  | 12.000 |
|   |                     |        |        |        |         | 1.2500  | 3.3333  | 7.5000 |
| 7 |                     |        |        |        |         |         | 2.0000  | 4.8889 |
|   |                     |        |        |        |         |         | 1.4104  | 4.0764 |
|   |                     |        |        |        |         |         | 1.5000  | 6.0000 |
|   |                     |        |        |        |         |         | 1.3333  | 4.0000 |

<sup>1,2,3,</sup> and <sup>4</sup> stand for the  $\sigma^{-2}$ MSRE of  $\hat{R}_{l,M}$ ,  $\hat{R}_{l,C}$ ,  $\hat{R}_{l}$  and  $\tilde{R}_{l}$ , respectively.

For n = 10 from Table 1, by empirical evidence, it is observed that in the sense of MSRE, BLIR is more accurate reconstructor than others. Furthermore, in the most cases the MSRE of  $\hat{R}_{l,M}$  is greater than others.

#### 5.2. Comparison based on PMC

Here, we use the PMC criterion to compare the reconstructors  $\hat{R}_l$ ,  $\hat{R}_l$ ,  $\hat{R}_{l,M}$  and  $\hat{R}_{l,C}$ . The measure is based on the probabilities of the relative closeness of competing estimators to an unknown parameter. First, we recall two definitions. If  $T_1$  and  $T_2$  are two estimators of a common parameter  $\theta$ , PMC of  $T_1$  relative to  $T_2$  is defined by

$$(5.8) \qquad PMC(T_1, T_2|\theta) = P(|T_1 - \theta| < |T_2 - \theta|), \ \forall \ \theta \in \Omega,$$

where  $\Omega$  is the parameter space. The estimator  $T_1$  is called Pitman–closer (with respect to  $\theta$ ) than  $T_2$  if and only if

$$PMC(T_1, T_2|\theta) \geq 1/2, \forall \theta \in \Omega,$$

with strict inequality holding for at least one  $\theta$ .

Based on a Type-II right censored sample  $(X_{1:n}, X_{2:n}, ..., X_{r:n})$  from  $\text{Exp}(0, \sigma)$  distribution, Balakrishnan *et al.* [8] compared BLUE and BLIE of  $\sigma$  in terms of PMC criterion. Also, Balakrishnan *et al.* [9] compared BLUP and BLIP of future order statistics under Type-II censoring for  $\text{Exp}(0, \sigma)$  distribution in terms of PMC criterion. Now, we compare each two reconstructors of  $R_l$  based on PMC criterion.

(i) PMC of  $\hat{R}_l$  relative to  $\tilde{R}_l$  (BLUR and BLIR): From (3.6), (4.3) and (5.8), we have

$$\begin{aligned} \pi_{\hat{R}_{l},\tilde{R}_{l}}(l,m;n) &= P(|\hat{R}_{l}-R_{l}| < |\tilde{R}_{l}-R_{l}|) \\ &= P\left(\left|R_{m+1} - \frac{m+1-l}{n-m-1}(R_{n}-R_{m+1}) - R_{l}\right| < \left|R_{m+1} - \frac{m+1-l}{n-m}\right| \\ &\times (R_{n}-R_{m+1}) - R_{l}\right| \right) \\ &= P\left(\frac{m+1-l}{(n-m-1)^{2}}(R_{n}-R_{m+1}) - \frac{2}{n-m-1}(R_{m+1}-R_{l}) < \frac{m+1-l}{(n-m)^{2}}\right) \\ &\times (R_{n}-R_{m+1}) - \frac{2}{n-m}(R_{m+1}-R_{l}) \right) \\ &= P\left(\frac{R_{n}-R_{m+1}}{R_{m+1}-R_{l}} < \frac{2}{(m+1-l)}\left(\frac{1}{n-m-1} + \frac{1}{n-m}\right)^{-1}\right). \end{aligned}$$

It can be shown that  $(R_n - R_{m+1})/(R_{m+1} - R_l)$  has the *F*-distribution with 2(n - m - 1) and 2(m + 1 - l) degrees of freedom. So, we express  $\pi_{\hat{R}_l, \widetilde{R}_l}(l, m; n)$  as

(5.9) 
$$\pi_{\hat{R}_l,\tilde{R}_l}(l,m;n) = F_{2(n-m-1),2(m+1-l)}\left(\frac{2(n-m)(n-m-1)}{(m+1-l)(2n-2m-1)}\right),$$

where  $F_{a,b}(x)$  stands for the cdf of *F*-distribution with *a* and *b* degrees of freedom at *x*. (ii) PMC of  $\hat{R}_l$  relative to  $\hat{R}_{l,M}$  (BLUR and MLR): From (3.6), (5.1) and (5.8), we have

$$\begin{aligned} \pi_{\hat{R}_l,\hat{R}_{l,M}}(l,m;n) &= P(|\hat{R}_l - R_l| < |\hat{R}_{l,M} - R_l|) \\ &= P(\left|R_{m+1} - \frac{m+1-l}{n-m-1}(R_n - R_{m+1}) - R_l\right| < \left|R_{m+1} - \frac{m-l}{n-m+1} \times (R_n - R_{m+1}) - R_l\right|) \end{aligned}$$

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(5.10) 
$$= F_{2(n-m-1),2(m+1-l)} \left( 2 \left( \frac{m+1-l}{n-m-1} + \frac{m-l}{n-m+1} \right)^{-1} \right).$$

(iii) PMC of  $\hat{R}_l$  relative to  $\hat{R}_{l,C}$  (BLUR and CMR): From (3.6), (5.3) and (5.8), we have

(5.11)  

$$\begin{aligned} \pi_{\hat{R}_{l},\hat{R}_{l,C}}(l,m;n) &= P(|\hat{R}_{l} - R_{l}| < |\hat{R}_{l,C} - R_{l}|) \\ &= P\left\{ \left( |R_{m+1} - \frac{m+1-l}{n-m-1}(R_{n} - R_{m+1}) - R_{l}| \right) \\ &< |R_{m+1} - \frac{m(1-M_{l,m})}{n-m}(R_{n} - R_{m+1}) - R_{l}| \right\} \\ &= F_{2(n-m-1),2(m+1-l)} \left( 2\left(\frac{m+1-l}{n-m-1} + \frac{m(1-M_{l,m})}{n-m}\right)^{-1} \right). \end{aligned}$$

(iv) PMC of  $\tilde{R}_l$  relative to  $\hat{R}_{l,M}$  (BLIR and MLR): Similarly, we have

$$\pi_{\widetilde{R}_{l},\widehat{R}_{l,M}}(l,m;n) = P(|\widetilde{R}_{l} - R_{l}| < |\widehat{R}_{l,M} - R_{l}|)$$

$$= P(|R_{m+1} - \frac{m+1-l}{n-m}(R_{n} - R_{m+1}) - R_{l}| < |R_{m+1} - \frac{m-l}{n-m+1} \times (R_{n} - R_{m+1}) - R_{l}|)$$

$$(5.12) = F_{2(n-m-1),2(m+1-l)}\left(\left(\frac{m+1-l}{n-m} + \frac{m-l}{n-m+1}\right)^{-1}\right).$$

(v) PMC of  $\tilde{R}_l$  relative to  $\hat{R}_{l,C}$  (BLIR and CMR):

(5.13) 
$$\pi_{\widetilde{R}_{l},\widehat{R}_{l,C}}(l,m;n) = F_{2(n-m-1),2(m+1-l)}\left(2\left(\frac{m+1-l}{n-m} + \frac{m(1-M_{l,m})}{n-m}\right)^{-1}\right).$$

(vi) PMC of  $\hat{R}_{l,M}$  relative to  $\hat{R}_{l,C}$  (MLR and CMR):

(5.14) 
$$\pi_{\hat{R}_{l,C},\hat{R}_{l,M}}(l,m;n) = F_{2(n-m-1),2(m+1-l)} \left( 2\left(\frac{m-l}{n-m+1} + \frac{m(1-M_{l,m})}{n-m}\right)^{-1} \right).$$

From the above results, one can examine the performance of any two reconstructors in terms of PMC criterion by comparing the values of median of the *F*-distribution with 2(n-m-1) and 2(m+1-l) degrees of freedom and corresponding values in (5.9)–(5.14), but, numerical computations are needed. We have computed the PMC probabilities in (5.9)– (5.14) for n = 5, 10, 15, 20 and all selected values of *l* and *m*. The results for n = 10 are presented in Tables 2–7, respectively, which examine how much a reconstructor is better than the other. The numerical results for n = 5, 15 and 20 are available with the authors, and will be sent on request to interested readers.

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Best Linear Unbiased and Invariant Reconstructors for the Past Records

|   |        |        |        |        | т      |        |        |        |
|---|--------|--------|--------|--------|--------|--------|--------|--------|
| l | 1      | 2      | 3      | 4      | 5      | 6      | 7      | 8      |
| 1 | 0.8894 | 0.8938 | 0.8219 | 0.6631 | 0.4423 | 0.2539 | 0.1554 | 0.1521 |
| 2 |        | 0.8758 | 0.8661 | 0.7603 | 0.5574 | 0.3337 | 0.1950 | 0.1713 |
| 3 |        |        | 0.8583 | 0.8271 | 0.6749 | 0.4377 | 0.2498 | 0.1960 |
| 4 |        |        |        | 0.8352 | 0.7703 | 0.5623 | 0.3267 | 0.2288 |
| 5 |        |        |        |        | 0.8034 | 0.6868 | 0.4332 | 0.2740 |
| 6 |        |        |        |        |        | 0.7570 | 0.5680 | 0.3393 |
| 7 |        |        |        |        |        |        | 0.6849 | 0.4375 |
| 8 |        |        |        |        |        |        |        | 0.5714 |

Table 2. Values of  $\pi_{\hat{R}_l,\tilde{R}_l}(l,m;n)$  for n = 10.

| Table 3. Values of $\pi_{\hat{R}_l,\hat{R}_{l,M}}(l,m;n)$ for $n = 10$ . |
|--|
|--|

|   |        |        |        |        | т      |        |        |        |
|---|--------|--------|--------|--------|--------|--------|--------|--------|
| l | 1      | 2      | 3      | 4      | 5      | 6      | 7      | 8      |
| 1 | 0.9396 | 0.9351 | 0.8812 | 0.7492 | 0.5346 | 0.3215 | 0.1946 | 0.1741 |
| 2 |        | 0.9314 | 0.9167 | 0.8345 | 0.6531 | 0.4154 | 0.2434 | 0.1965 |
| 3 |        |        | 0.9206 | 0.8898 | 0.7646 | 0.5318 | 0.3104 | 0.2254 |
| 4 |        |        |        | 0.9057 | 0.8482 | 0.6620 | 0.4023 | 0.2638 |
| 5 |        |        |        |        | 0.8842 | 0.7811 | 0.5248 | 0.3170 |
| 6 |        |        |        |        |        | 0.8503 | 0.6700 | 0.3942 |
| 7 |        |        |        |        |        |        | 0.7901 | 0.5100 |
| 8 |        |        |        |        |        |        |        | 0.6667 |

Table 4. Values of  $\pi_{\hat{R}_l,\hat{R}_{l,C}}(l,m;n)$  for n = 10.

|   |        |        |        |        | т      |        |        |        |
|---|--------|--------|--------|--------|--------|--------|--------|--------|
| l | 1      | 2      | 3      | 4      | 5      | 6      | 7      | 8      |
| 1 | 0.9141 | 0.9097 | 0.8519 | 0.7012 | 0.4763 | 0.2739 | 0.1646 | 0.1560 |
| 2 |        | 0.8758 | 0.8893 | 0.7919 | 0.5924 | 0.3581 | 0.2064 | 0.1759 |
| 3 |        |        | 0.8814 | 0.8526 | 0.7081 | 0.4666 | 0.2646 | 0.2016 |
| 4 |        |        |        | 0.8601 | 0.7991 | 0.5943 | 0.3461 | 0.2358 |
| 5 |        |        |        |        | 0.8308 | 0.7184 | 0.4583 | 0.2832 |
| 6 |        |        |        |        |        | 0.7876 | 0.5985 | 0.3521 |
| 7 |        |        |        |        |        |        | 0.7182 | 0.4566 |
| 8 |        |        |        |        |        |        |        | 0.6002 |

| Table 5. Values of $\pi_{\widetilde{R}_l, \hat{R}_{l,M}}(l, m; n)$ for $n = 10$ . |  |
|---|--|
|   |  |

|   |        |        |        | $\mathbf{n}_{l},\mathbf{n}_{l},$ | 141    |        |        |        |
|---|--------|--------|--------|----------------------------------|--------|--------|--------|--------|
|   |        |        |        |                                  | m      |        |        |        |
| l | 1      | 2      | 3      | 4                                | 5      | 6      | 7      | 8      |
| 1 | 0.9461 | 0.9447 | 0.9030 | 0.7974                           | 0.6142 | 0.4118 | 0.2786 | 0.2663 |
| 2 |        | 0.9397 | 0.9305 | 0.8675                           | 0.7218 | 0.5116 | 0.3402 | 0.2985 |
| 3 |        |        | 0.9315 | 0.9104                           | 0.8157 | 0.6255 | 0.4199 | 0.3390 |
| 4 |        |        |        | 0.9207                           | 0.8808 | 0.7410 | 0.5212 | 0.3912 |
| 5 |        |        |        |                                  | 0.9060 | 0.8356 | 0.6421 | 0.4602 |
| 6 |        |        |        |                                  |        | 0.8847 | 0.7658 | 0.5528 |
| 7 |        |        |        |                                  |        |        | 0.8521 | 0.6735 |
| 8 |        |        |        |                                  |        |        |        | 0.8000 |

|   |        |        |        |        | т      |        |        |        |
|---|--------|--------|--------|--------|--------|--------|--------|--------|
| l | 1      | 2      | 3      | 4      | 5      | 6      | 7      | 8      |
| 1 | 0.9204 | 0.9204 | 0.8752 | 0.7496 | 0.5487 | 0.3482 | 0.2290 | 0.2264 |
| 2 |        | 0.8835 | 0.9041 | 0.8264 | 0.6577 | 0.4394 | 0.2813 | 0.2536 |
| 3 |        |        | 0.8918 | 0.8745 | 0.7589 | 0.5491 | 0.3508 | 0.2882 |
| 4 |        |        |        | 0.8741 | 0.8326 | 0.6675 | 0.4422 | 0.3331 |
| 5 |        |        |        |        | 0.8509 | 0.7714 | 0.5572 | 0.3931 |
| 6 |        |        |        |        |        | 0.8187 | 0.6836 | 0.4750 |
| 7 |        |        |        |        |        |        | 0.7717 | 0.5860 |
| 8 |        |        |        |        |        |        |        | 0.7062 |

Table 6. Values of  $\pi_{\tilde{R}_{l},\hat{R}_{lc}}(l,m;n)$  for n = 10.

Table 7. Values of  $\pi_{\hat{R}_{l,C},\hat{R}_{l,M}}(l,m;n)$  for n = 10.

|   |        |        |        |        | т      |        |        |        |
|---|--------|--------|--------|--------|--------|--------|--------|--------|
| l | 1      | 2      | 3      | 4      | 5      | 6      | 7      | 8      |
| 1 | 0.9727 | 0.9576 | 0.9278 | 0.8348 | 0.6584 | 0.4475 | 0.3002 | 0.2786 |
| 2 |        | 0.9397 | 0.9498 | 0.8957 | 0.7618 | 0.5508 | 0.3659 | 0.3125 |
| 3 |        |        | 0.9569 | 0.9326 | 0.8489 | 0.6663 | 0.4509 | 0.3557 |
| 4 |        |        |        | 0.9486 | 0.9077 | 0.7796 | 0.5578 | 0.4118 |
| 5 |        |        |        |        | 0.9378 | 0.8691 | 0.6829 | 0.4862 |
| 6 |        |        |        |        |        | 0.9224 | 0.8070 | 0.5863 |
| 7 |        |        |        |        |        |        | 0.8985 | 0.7167 |
| 8 |        |        |        |        |        |        |        | 0.8576 |

From our numerical results and by empirical evidence, we observe that:

(1) The BLIR (CMR) is Pitman-closer than BLUR (BLUR) for the following situations:

(i) when n = 10, 20, m = n/2 + a and l = 1, 2, ..., 2a + 1, (a = 0, 1, ..., n/2 - 2);(ii) when n = 5, 15, m = (n + 1)/2 + a and l = 1, 2, ..., 2a + 2, (a = 0, 1, ..., (n + 1)/2 - 3);otherwise the BLUE (BLUE) is Pitman-closer than the BLE (CMR) specially for

otherwise, the BLUR (BLUR) is Pitman-closer than the BLIR (CMR) specially for l = m. [See for example Table 2 (Table 4) when n = 10].

(2) The MLR (CMR) is Pitman-closer than BLUR (BLIR) in the following situations:
(i) when n = 10,20, m = n/2 + a and l = 1,2,...,2a, (a = 1,...,n/2-2);
(ii) when n = 5,15, m = (n + 1)/2 + a and l = 1,2,...,2a + 1, (a = 0,1,...,(n + 1)/2 - 3);
otherwise, the BLUR (BLIR) is Pitman-closer than the MLR (CMR) specially for

l = m. [See for example Table 3 (Table 6) when n = 10].

(3) The MLR (MLR) is Pitman-closer than BLIR (CMR) for the following situations:
(i) when n = 10,20, m = n/2 + a and l = 1,2,...,2a - 1, (a = 1,...,n/2 - 2);
(ii) when n = 5,15, m = (n+1)/2 + a and l = 1,2,...,2a, (a = 1,...,(n+1)/2 - 3); otherwise, the BLIR (CMR) is Pitman-closer than the MLR (MLR) specially for l = m. [See for example Table 5 (Table 7) when n = 10].

## 6. A real example

To illustrate the performance of the proposed reconstructors, we use a real data set concerning the times (in minutes) between 48 consecutive telephone calls to a company's switchboard which is presented by Castillo *et al.* [11] where the authors found that the exponential distribution gave an adequate fit. Table 8 contains the corresponding data.

| Table 0 | . Thice | , (iii iiii | nutes) i |      | 1 40 00 | isceuti | ve cans. |
|---------|---------|-------------|----------|------|---------|---------|----------|
| 1.34    | 0.14    | 0.33        | 1.68     | 1.86 | 1.31    | 0.83    | 0.33     |
| 2.20    | 0.62    | 3.20        | 1.38     | 0.96 | 0.28    | 0.44    | 0.59     |
| 0.25    | 0.51    | 1.61        | 1.85     | 0.47 | 0.41    | 1.46    | 0.09     |
| 2.18    | 0.07    | 0.02        | 0.64     | 0.28 | 0.68    | 1.07    | 3.25     |
| 0.59    | 2.39    | 0.27        | 0.34     | 2.18 | 0.41    | 1.08    | 0.57     |
| 0.35    | 0.69    | 0.25        | 0.57     | 1.90 | 0.56    | 0.09    | 0.28     |

Table 8. Times (in minutes) between 48 consecutive calls.

From the data of Table 8, six upper records extracted which are 1.34, 1.68, 1.86, 2.20, 3.20 and 3.25. Assuming that only  $R_{m+1}, ..., R_6$  have been observed, we show that how one can reconstruct the missing record values and compare the results with the exact values. With this in mind, we compute different point reconstructors presented in the paper for the lost record value  $R_l$  ( $1 \le l \le m$ ). Using (3.6), (4.3), (5.1) and (5.3) the values of the BLUR, BLIR, MLR and CMR are obtained, respectively. The results are presented in Table 9 for m = 3 and 4 and  $1 \le l \le m$ .

Table 9. The numerical values of point reconstructors.

| $\hat{R}_{l,C}$ |
|-----------------|
| .366            |
| .675            |
| .983            |
|                 |
| 3.116           |
| 3.139           |
| 3.161           |
| 3.184           |
| 3               |

For m = 3, the observed records are  $R_4 = 2.20$ ,  $R_5 = 3.20$  and  $R_6 = 3.25$ . Based on these observations we have reconstructed the first three records. From Table 9, it is observed that for m = 3 the point reconstructors are reasonably close to the exact values. When m = 4, it is assumed that two records  $R_5 = 3.20$  and  $R_6 = 3.25$  are observed. In this case, the point reconstructors are not close to the exact values but not too far. It should be mentiond that one example does not tell us much more.

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